

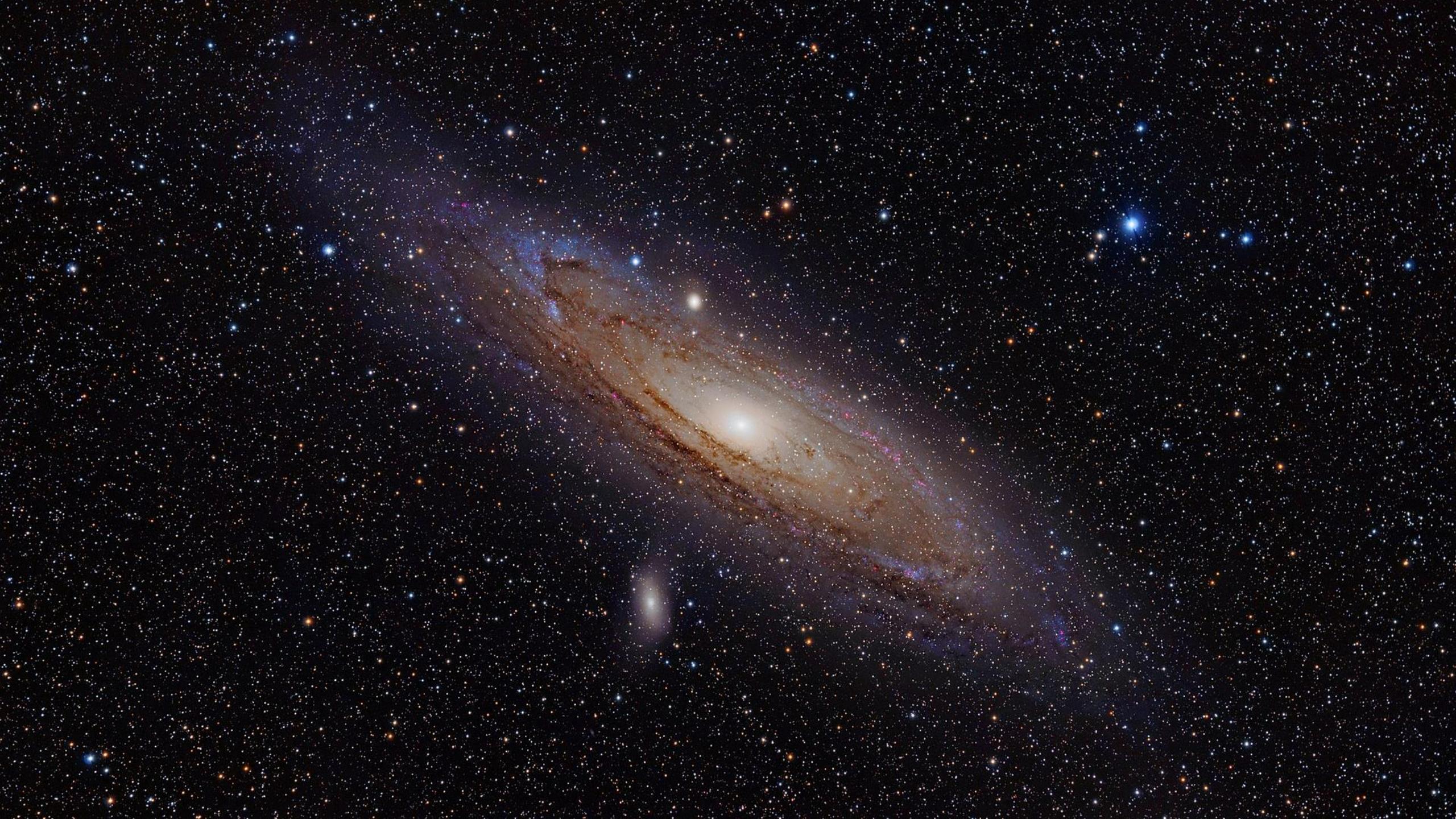
# Precision searches for new physics using optically levitated sensors

Gadi Afek

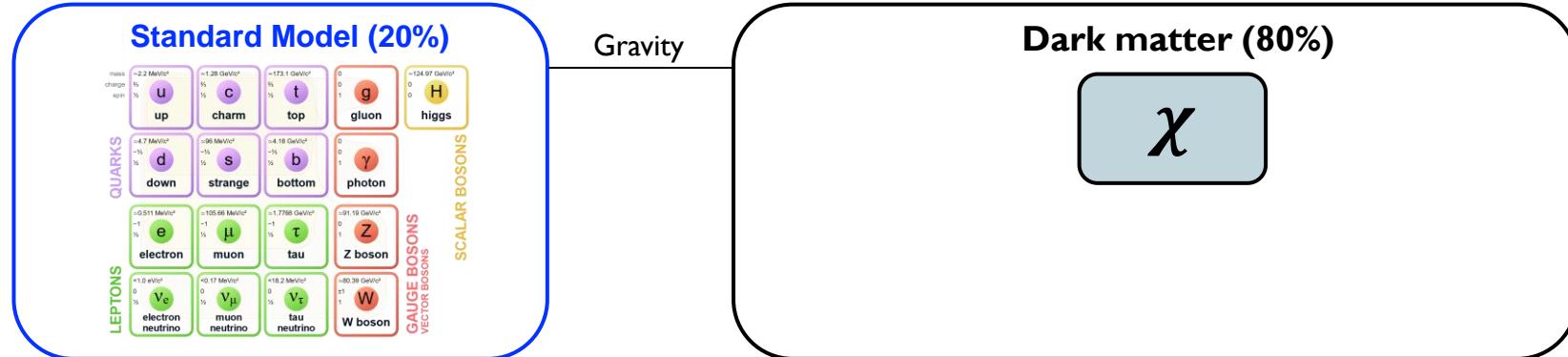
Yale University

Quantum Science Seminar  
hot topics session, July 2021

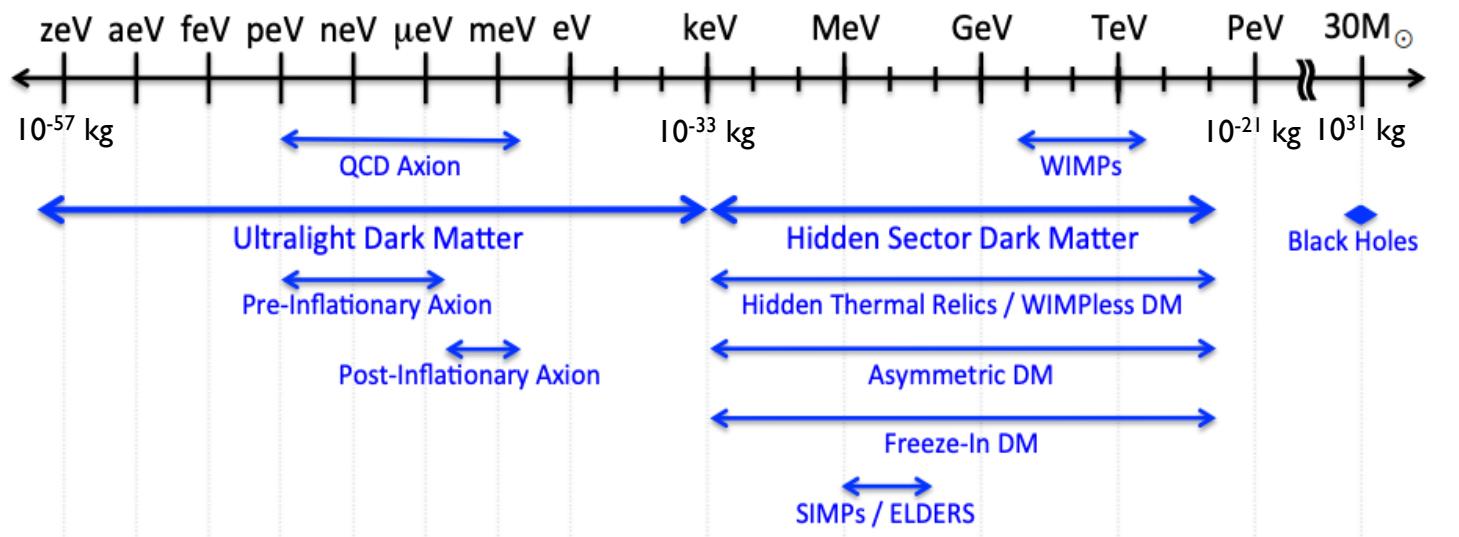




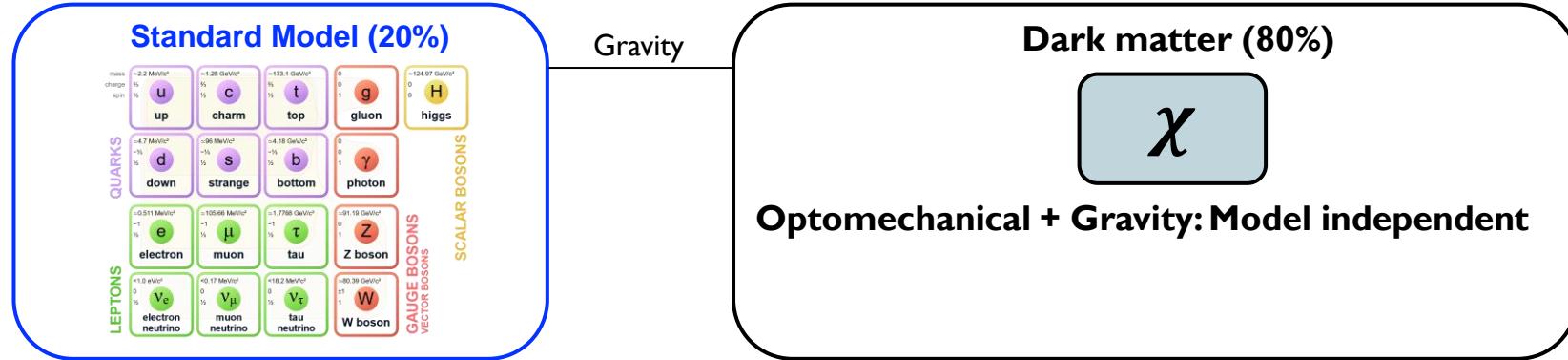
# WHAT DO WE KNOW SO FAR?



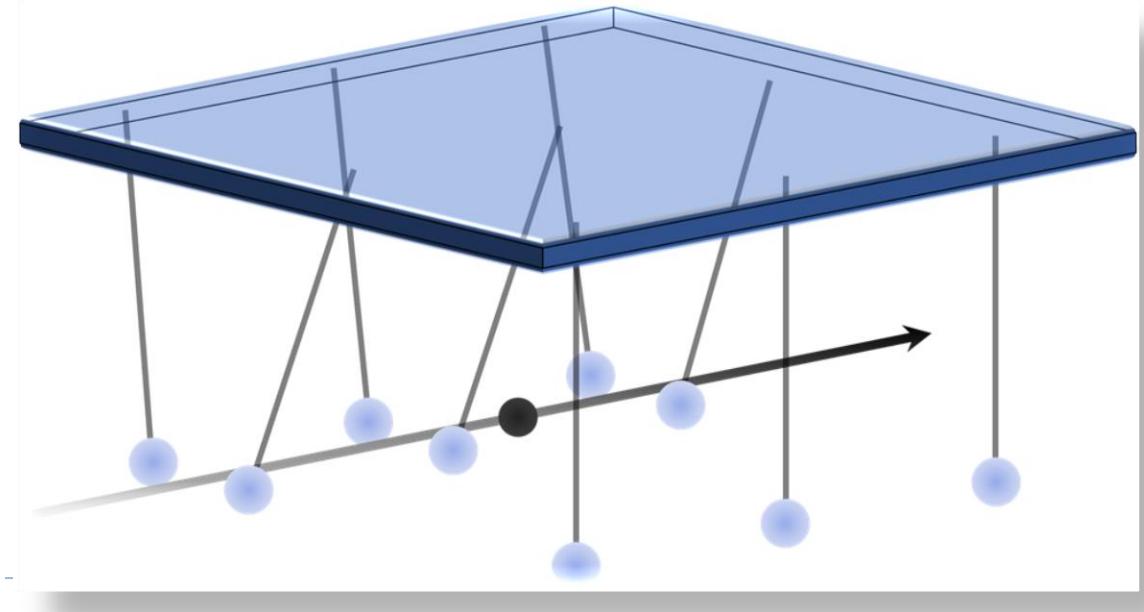
- ❑ Vast majority of terrestrial searches focus on **WIMPs** and **axions**
  - ❑ Compelling + solve other problems
  - ❑ Ton-scale WIMP searches haven't seen anything yet..
  - ❑ Axion parameter space getting bigger (quantum!)
- ❑ Many other models allowed



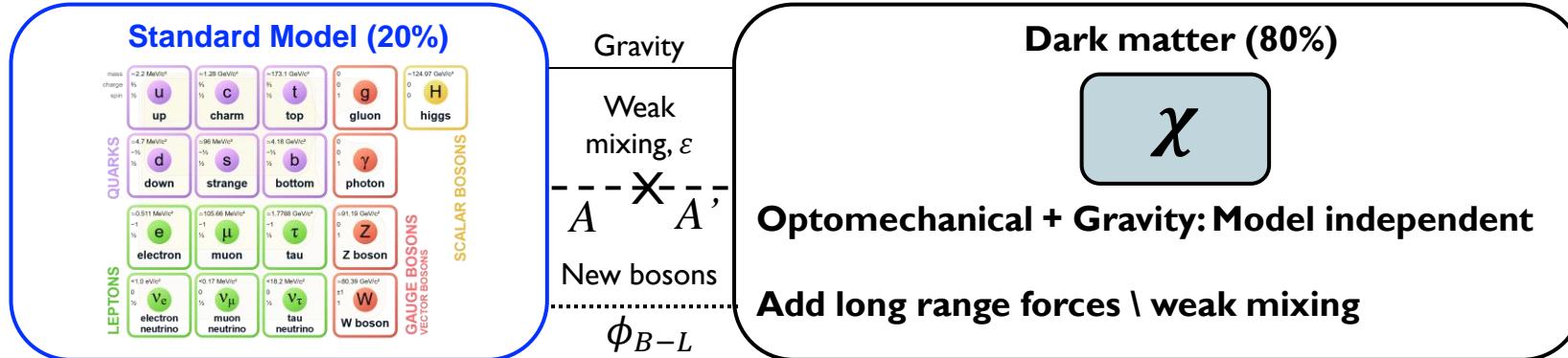
# MODEL-INDEPENDENT SEARCHES?



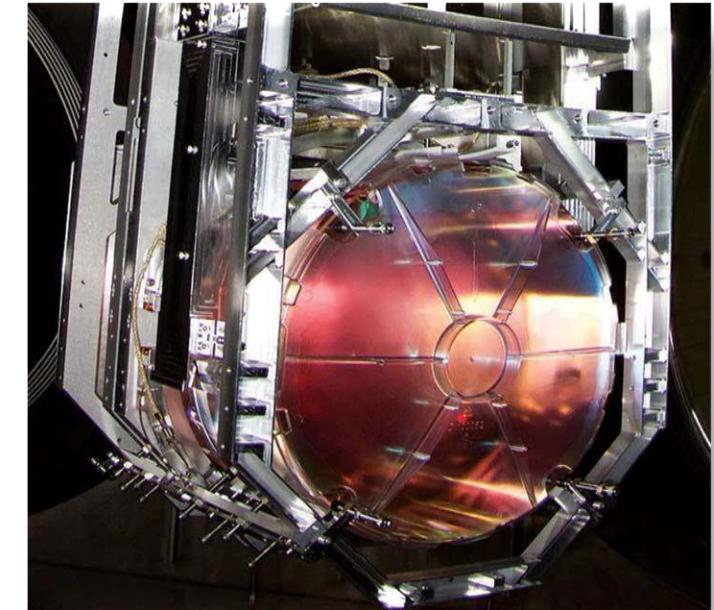
- What would a **model-independent** search look like?
- Search for **gravitationally induced** kicks from a passing dark matter particle in an **array of test masses**
- For DM masses  $> m_{Pl}$  might be possible but **very ambitious** (30dB beyond SQL, cryogenic environment,  **$10^9$  sensors**...)



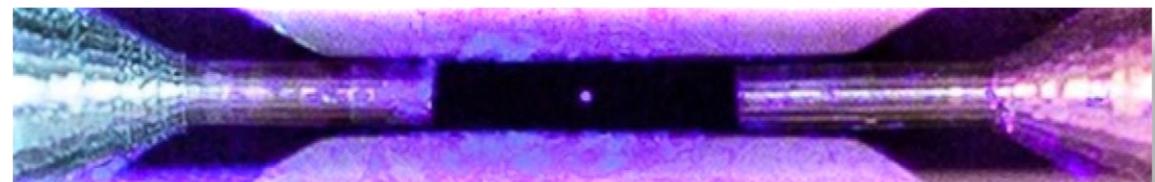
# OPTOMECHANICAL SENSORS



- Opto-mechanical systems are **VERY precise force sensors**
- Control and measurement of **large range of test masses** (from  $10^{-21} \text{ gr}$  to  $10^3 \text{ gr}$ )
- We use  $\sim 10^{-8} \text{ gr}$  microspheres with  **$\sim 100 \text{ ng}/\text{Hz}^{1/2}$  acceleration sensitivity (technically limited)**

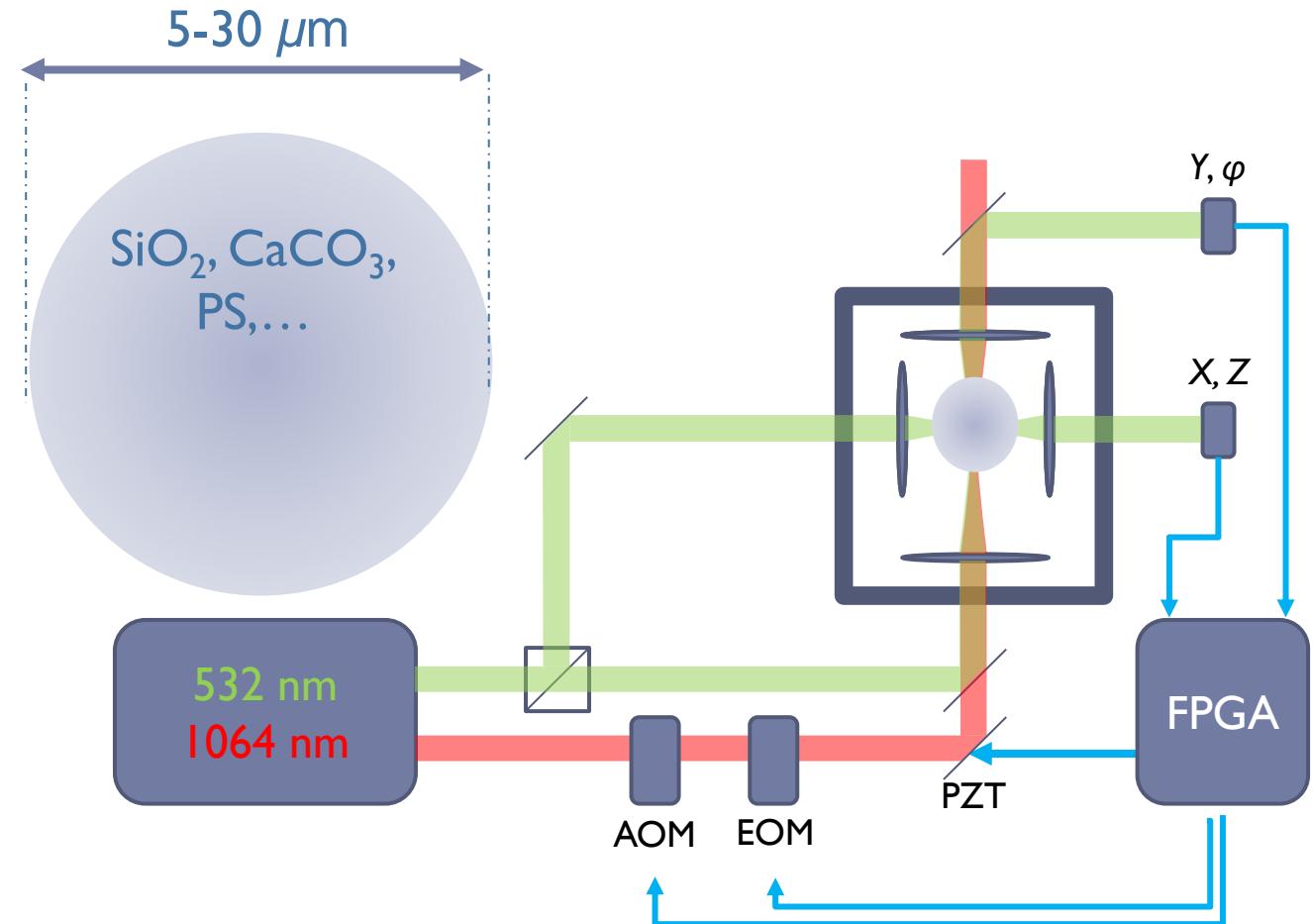


<https://www.newscientist.com/article/2084742-gravitational-wave-hunters-gear-up-to-detect-extreme-black-holes/>



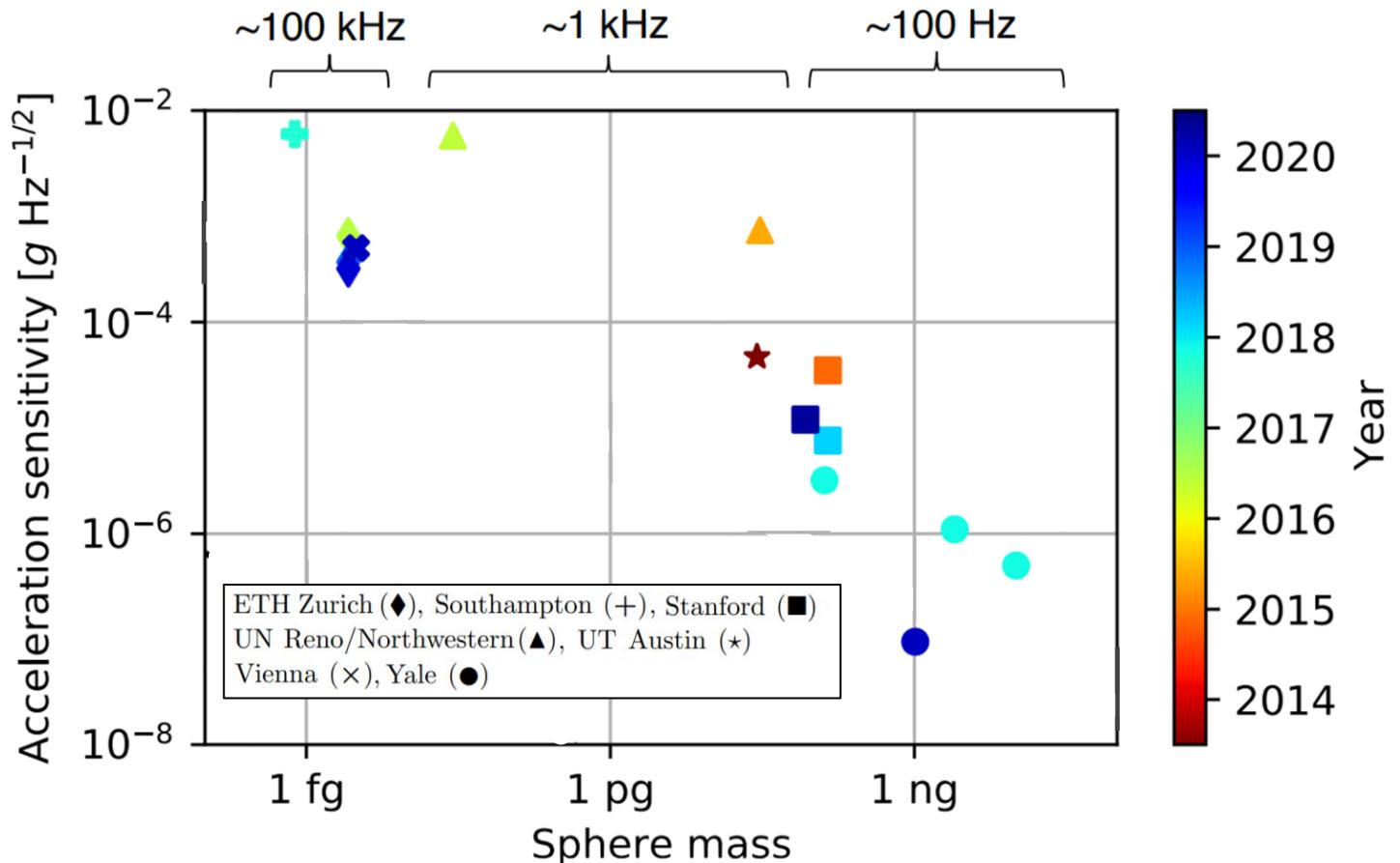
# THE SYSTEM

- Variety of materials and sizes, isolated electrically and thermally
- Large spheres → better acceleration sensitivity, DM searches couple to # constituents in sensor
- Low NA gravito-optical configuration →  $\sim \mu\text{m}$  probing distances
- Trap > 1 month → LONG integration times
- Nanospheres already at SQL.  $\mu$ spheres still ~100 above (technically limited)



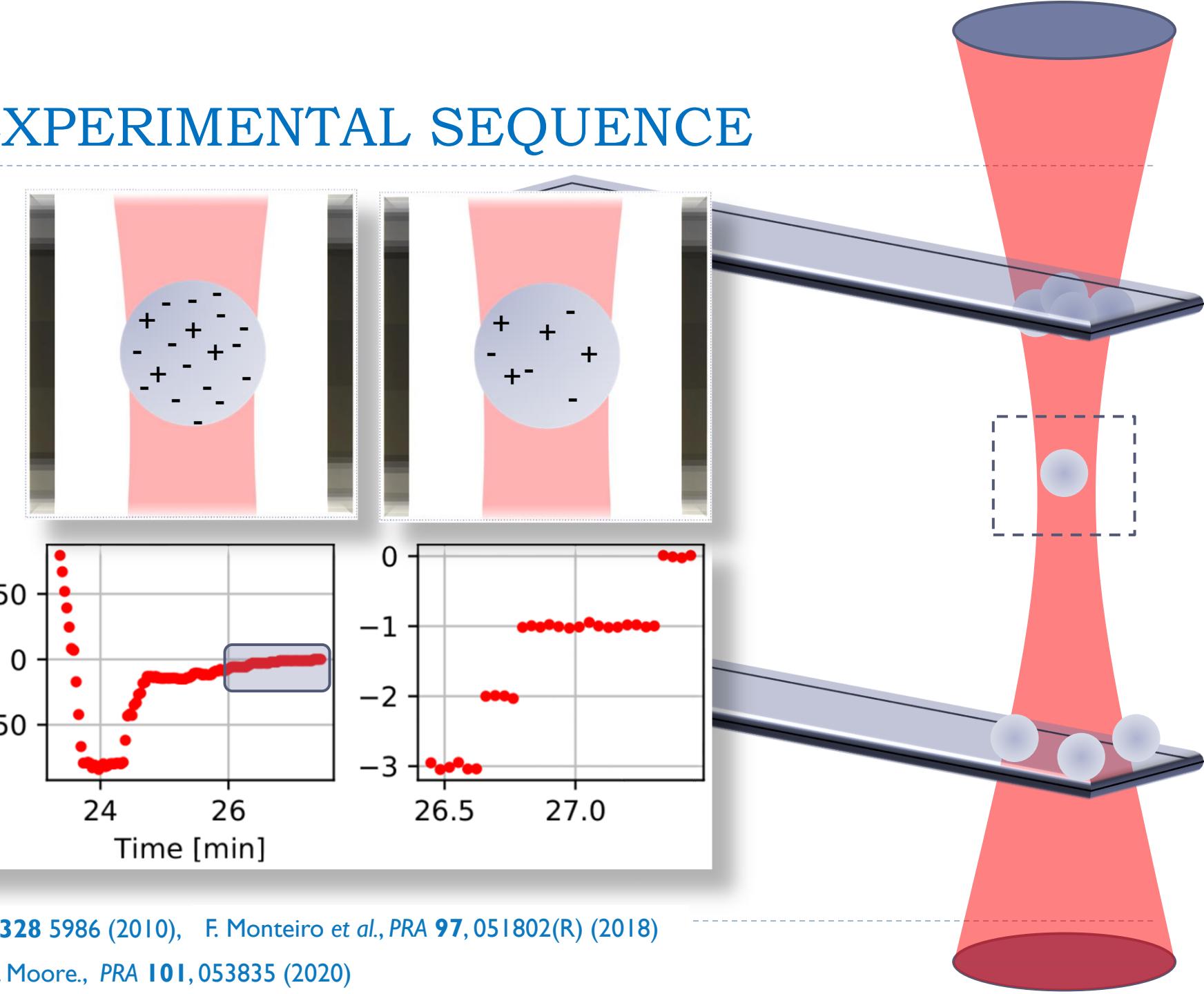
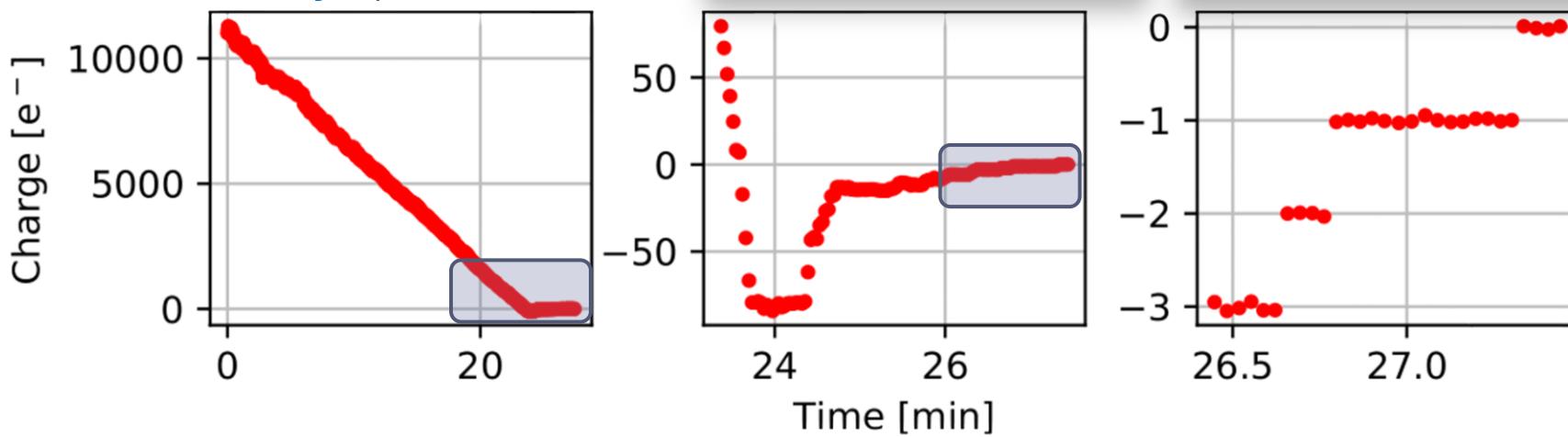
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# EXPERIMENTAL SEQUENCE

- Trap at high pressure
- pump down to  $\sim 10^{-7}$  torr
- Discharge with **single e precision**  
(Charging rates  $< 1$  e/week  $\sim 1$  yA)

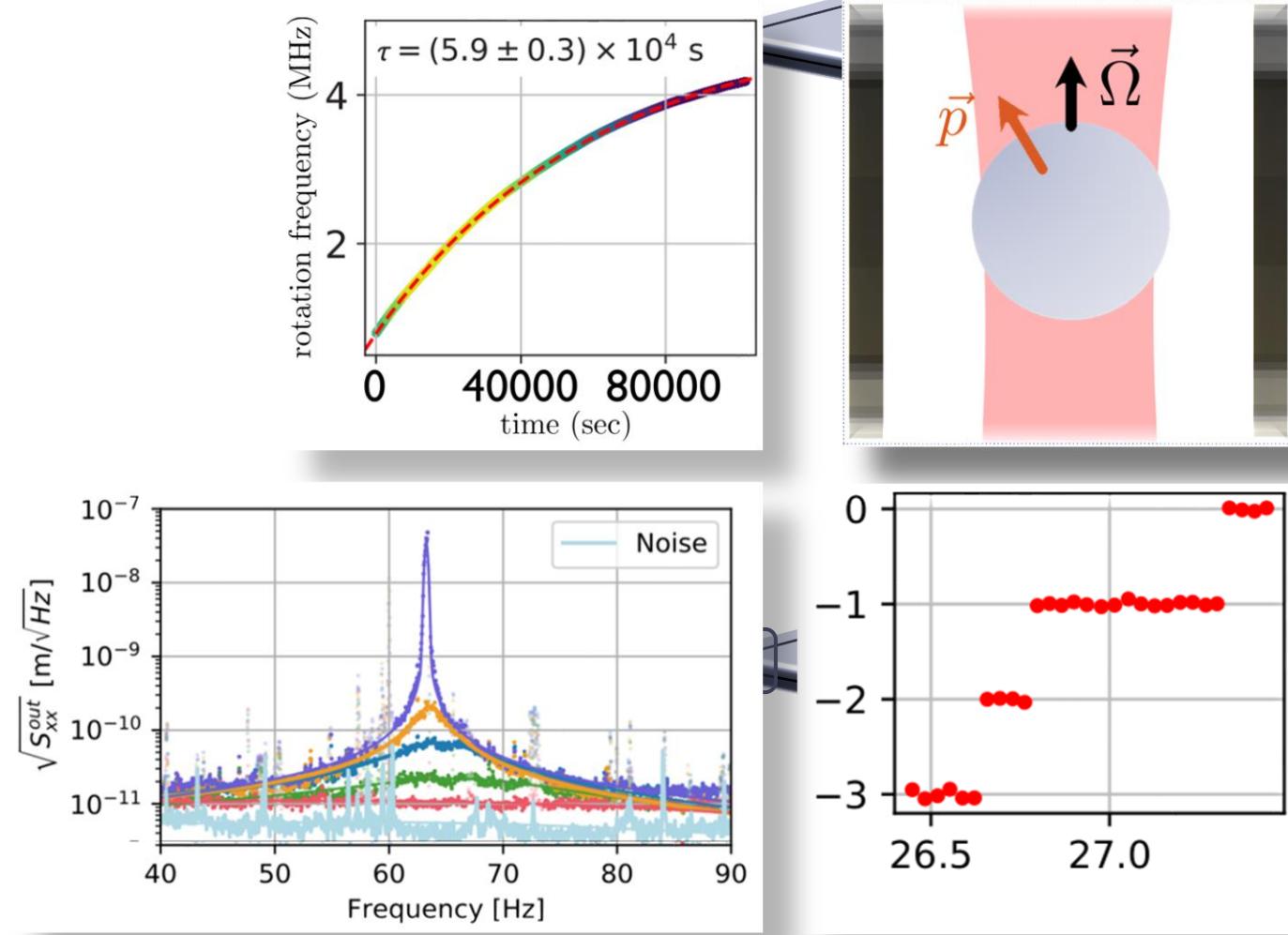


T. Li, S. Kheifets, D. Medellin, M. G. Raizen, *Science* **328** 5986 (2010), F. Monteiro et al., *PRA* **97**, 051802(R) (2018)

F. Monteiro, W. Li, **GA**, C.L. Li, M. Mossman and D. Moore., *PRA* **101**, 053835 (2020)

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- Spin up **to  $\sim 10$  MHz**
- **Cool COM to  $\sim 50 \mu\text{K}$**



T. Li, S. Kheifets, D. Medellin, M. G. Raizen, *Science* **328** 5986 (2010), F. Monteiro *et al.*, *PRA* **97**, 051802(R) (2018)

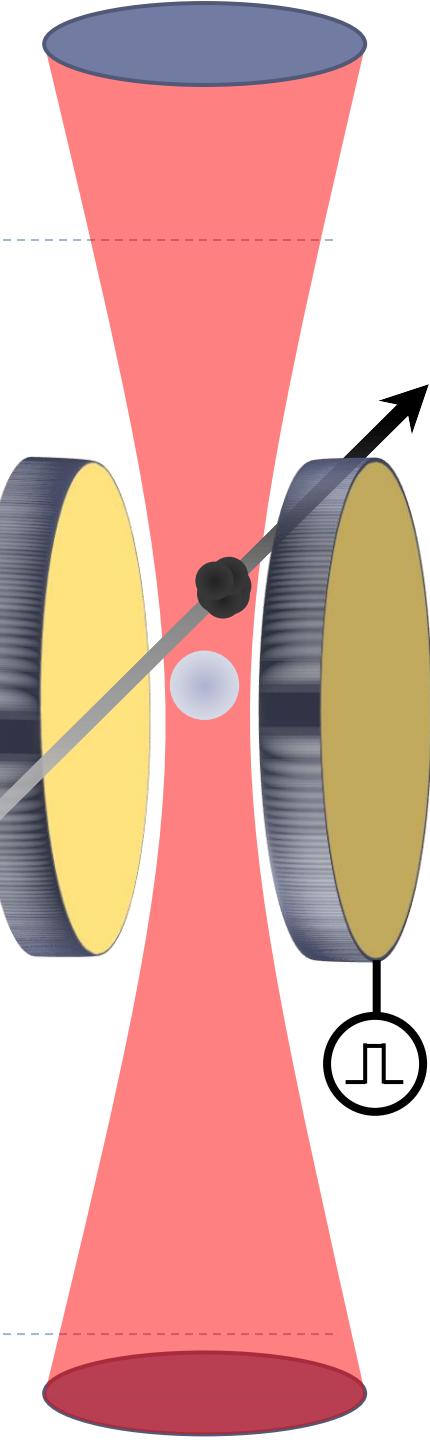
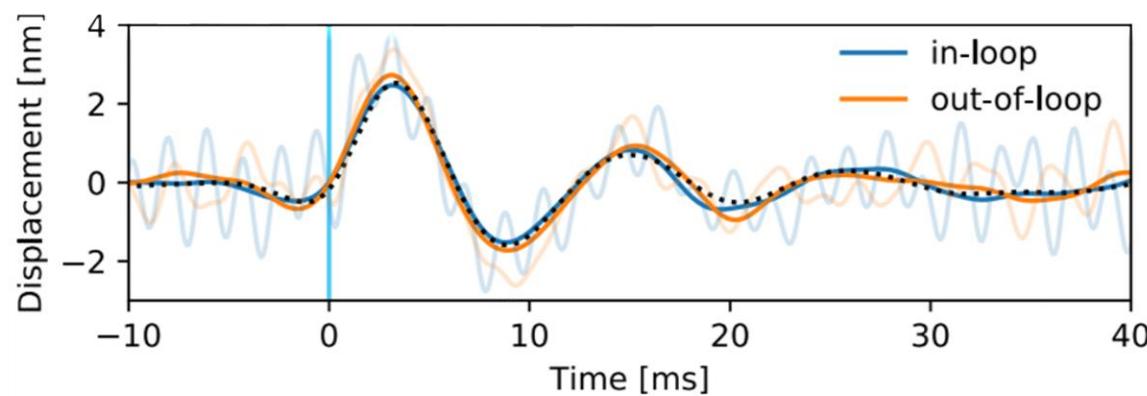
F. Monteiro, W. Li, **GA**, C.L. Li, M. Mossman and D. Moore., *PRA* **101**, 053835 (2020)

# DM-INDUCED RECOILS

- Consider heavy **DM** particles
- Interaction mediated by a **long-range** force carrier  $m_\phi \lesssim \text{eV}$
- **Coherent enhancement!**
- **Low momentum threshold**  
 $\sim 200 \text{ MeV}/c \sim 0.1 \text{ aN sec}$
- Total exposure: **5 ( $10^{-9} \text{ gr X day}$ )**

$$V(r) = \alpha_n N_n \frac{e^{-m_\phi r}}{r}$$

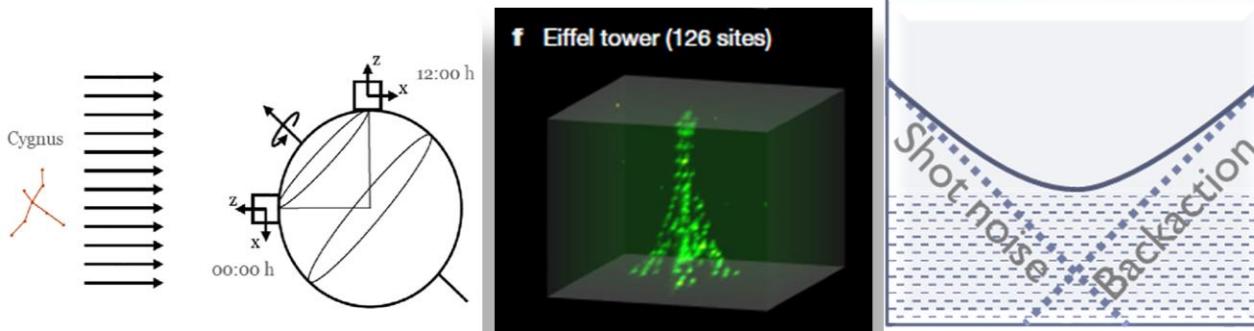
$$\sigma \sim N_n^2 \sim 10^{29}$$



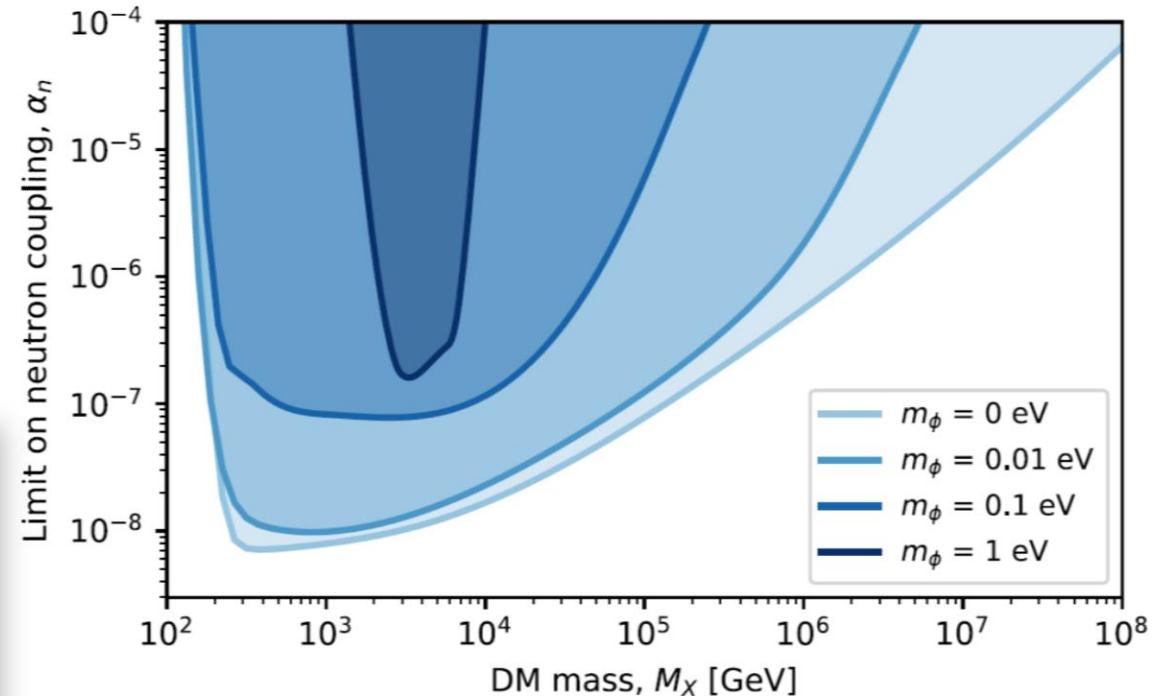
# MODEL-INDEPENDENT LIMITS

$$V(r) = \alpha_n N_n \frac{e^{-m_\phi r}}{r}$$

- ❑ Assuming specific composite dark matter model, can **compare to WIMP detectors**
- ❑ For sufficiently light mediators and large composite particles, **many orders-of-magnitude more sensitive**
- ❑ Only **first proof-of-principle. What's next?**



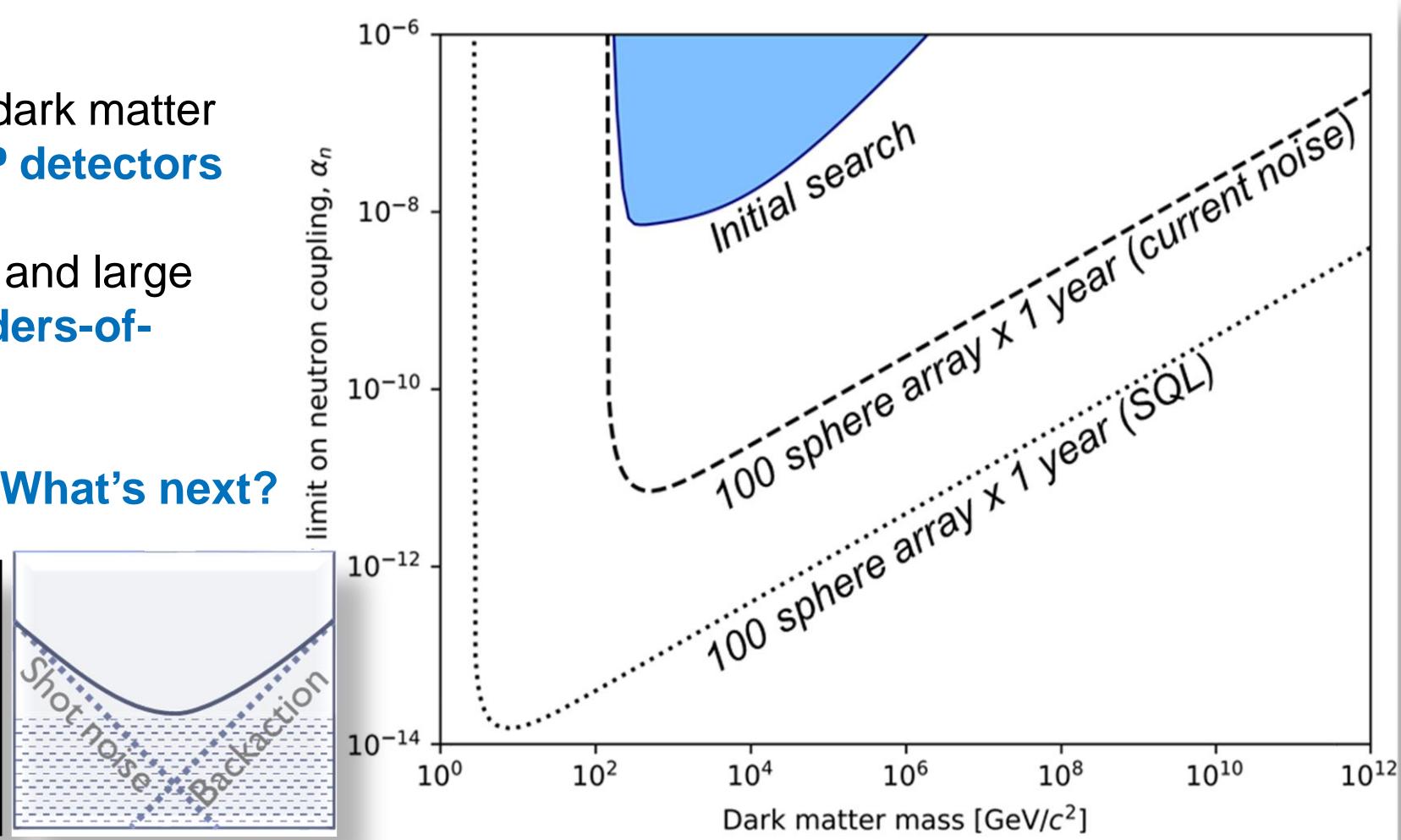
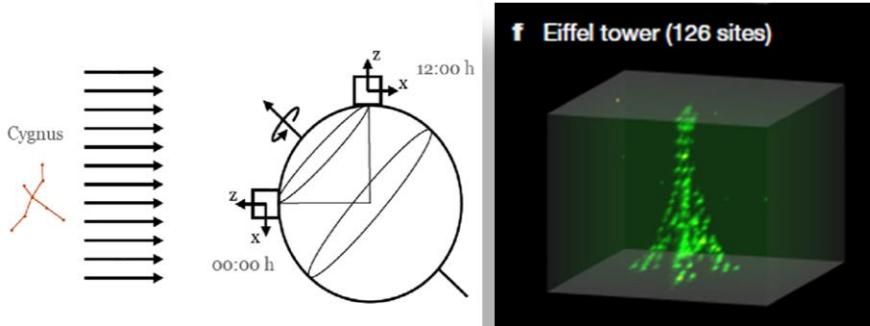
XENON1T, LUX, SuperCDMS, ... (> ton year)



# MODEL-INDEPENDENT LIMITS

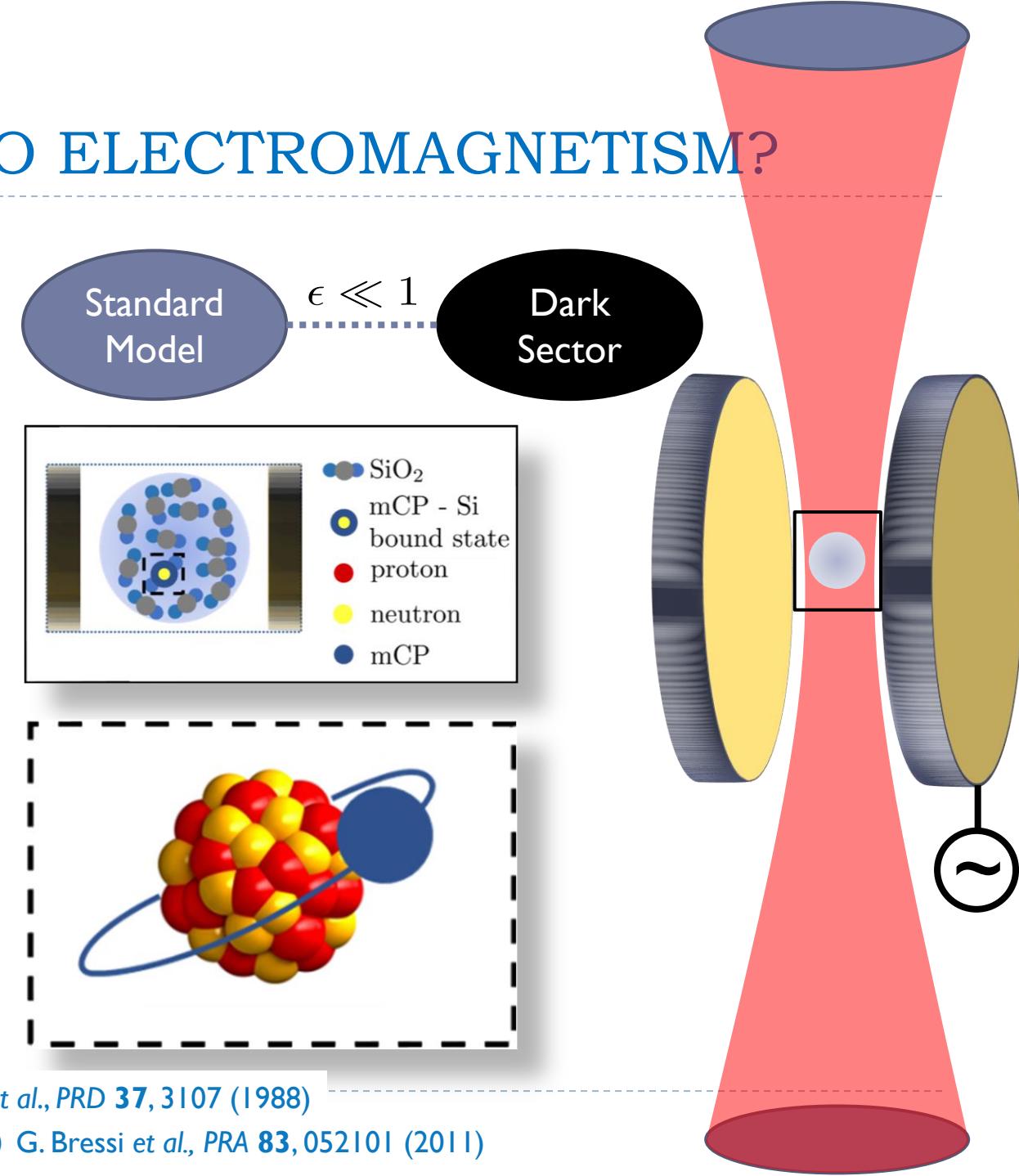
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# WHAT IF DM COUPLES TO ELECTROMAGNETISM?

- ❑ Particles with unity charge under new dark force can have fractional charge under electromagnetism
- ❑ Charge/mass ratio  $\sim 10^9$  worse than, e.g., single  $\text{Sr}^{+2}$  ion or  $\sim 10^{14}$  worse than a single  $e^-$
- ❑ Protons and electrons form bound states
- ❑ The experiment: Calibrate response function to known charge, discharge, spin up apply high voltage, measure correlation
- ❑ Neutrality of matter



D. Carney, H. Haffner, D. Moore and J. Taylor *arXiv:2104.05737* (2021) J. Baumann et al., *PRD* **37**, 3107 (1988)

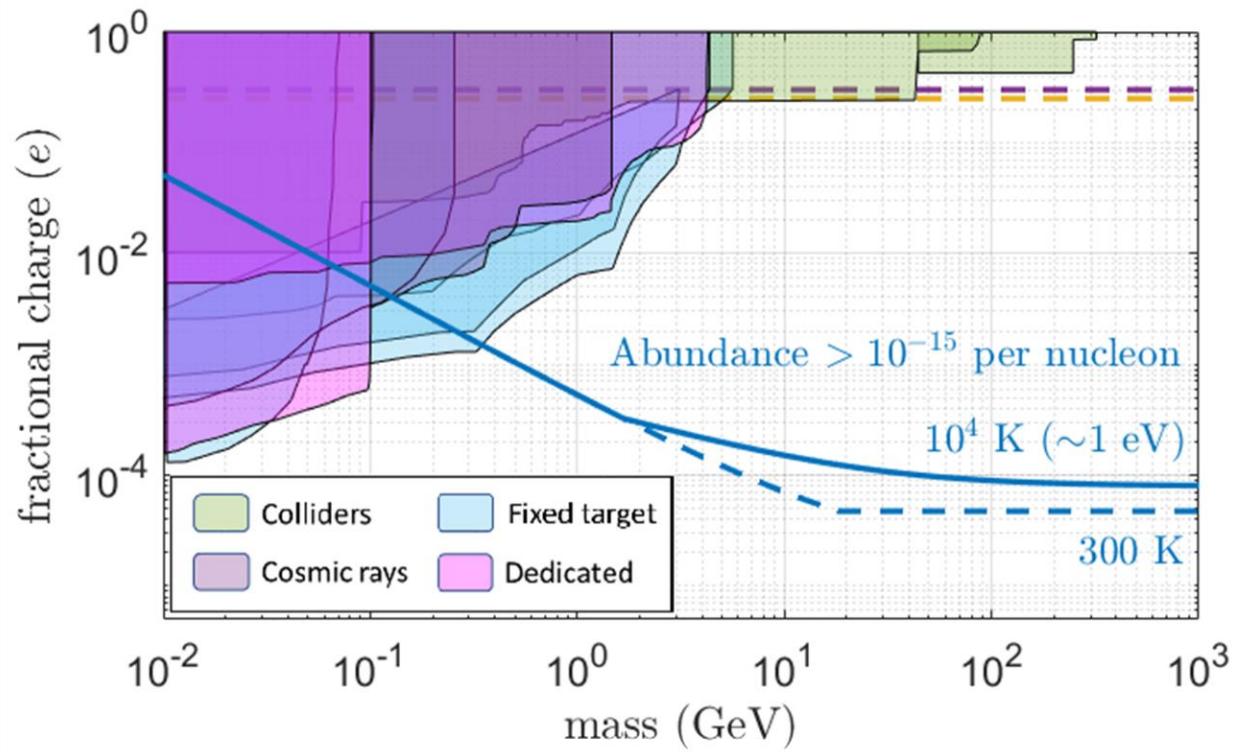
GA, F. Monteiro, J. Wang, B. Siegel, S. Ghosh and D. Moore, *PRD* **104**, 012004 (2021) G. Bressi et al., *PRA* **83**, 052101 (2011)

# LIMITS ON MILlichARGED PARTICLES BOUND TO NORMAL MATTER

- Total mass  $\sim 76 \times 10^{-9}$  gr



- Probe **deep into  $10^{-17}$  / nucleon**, << abundance of naturally occurring stable elements (10 parts in a quintillion...)
- Holds even in **comparison** with ambitious **future experiment projections**
- **$10^{-19}$  e / nucleon limit <  $10^{-21}$  e best**

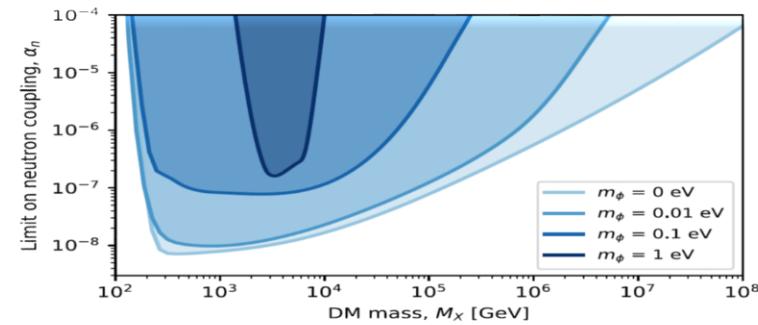


GA, F. Monteiro, J. Wang, B. Siegel, S. Ghosh and D. Moore, PRD **104**, 012004 (2021)

M. Marinelli et al., PhysRep **85** 161 (1982), P.C. Kim et al, PRL **99** 161804 (2007), D. C. Moore et al., PRL **113**, 251801 (2014), M. Pospelov et al., arXiv:2012.03957 (2020)

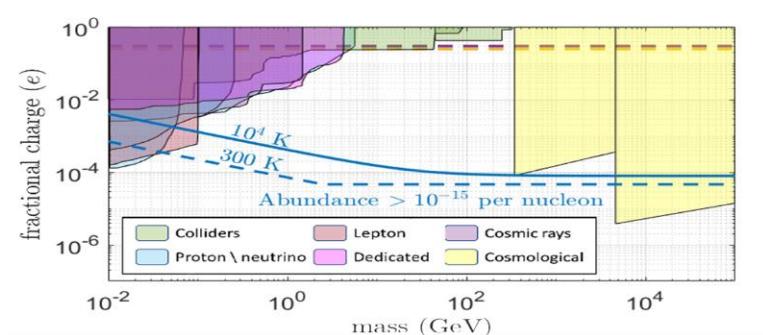
# WHAT'S NEXT, THEN?

## Search for **recoils** from composite DM



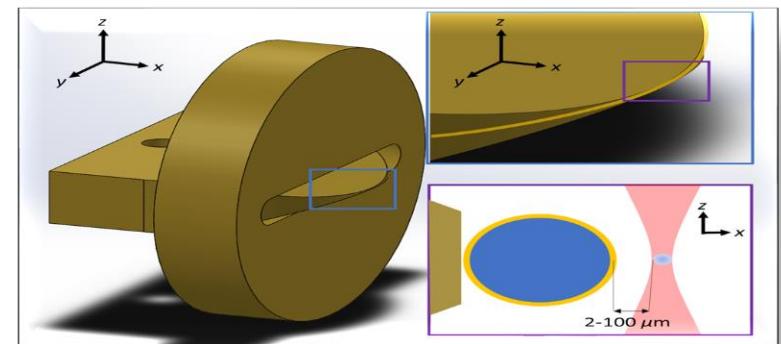
F. Monteiro, **GA**, D. Carney, G. Krnjaic, J. Wang and D. Moore., PRL 125, 181102 (2020)

## Testing **charge quantization** and search for mCP

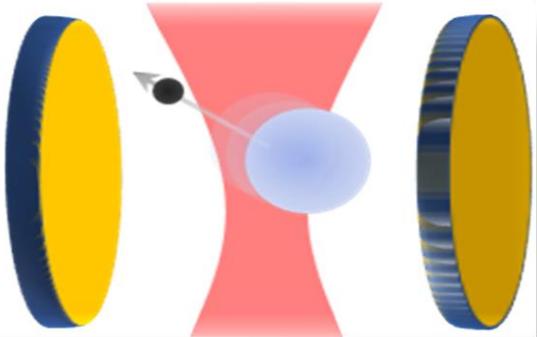


**GA**, F. Monteiro, J. Wang, B. Siegel, S. Ghosh and D. Moore., PRD 104, 012004 (2021)

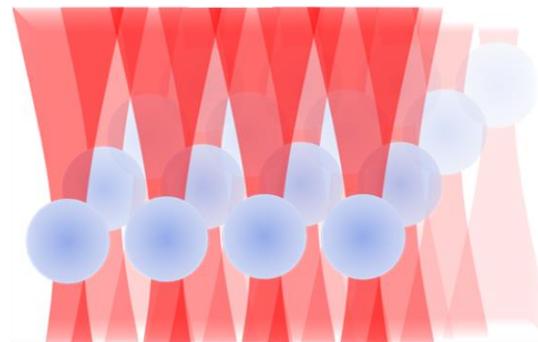
## Testing **Newton's law** at $\sim \mu\text{m}$ distances



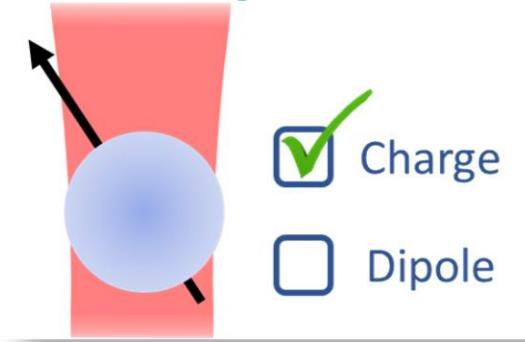
## Nuclear forensics from $\alpha/\beta$ decays



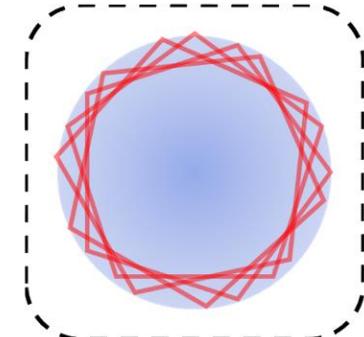
## Arrays of ng masses



## Controlling **dipole** backgrounds



## WGM spectroscopy



# THANK YOU!

